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PETROLOGICAL AND COSMOCHEMICAL EVIDENCE FOR VARIABLE C/O RATIO DURING FORMATION OF EL3 CHONDRITES OF ALMAHATA SITTA TC₃ ASTEROID.

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Primitive EL3 chondrites in Almahata Sitta asteroid (MS-17 & MS-177) contain two distinct lithologies: CaS-rich chondrules with graphite books and heterogeneous FeNi-metal nodules in matrix. FeNi-nodules consist of conglomerates of concentrically zoned individual C-bearing FeNi-pebbles (20-50µm in diameter). They enclose FeS, (FeNi)₃P, Fe₃C, Sinoite (**Si₂N₂O**), enstatite, graphite flames and no quenched immiscible liquids hence negating claimed impact or preaccretionary melting. We report a novel adouble track condensation sequence in separate FeNi-nodules: (1) **CaS⇒sinoite⇒enstatite**. Here enstatite largely consumed sinoite and (2) **CaS⇒sinoite⇒graphite flames** latter track evidencing abrupt increase in the C/O ratio (0.83-1.03) of the solar gas after sinoite. REE patterns of CaS in chondrules and in FeNi nodules are flat with negative Eu anomaly indicative of condensation from an REE fractionated source. δ¹³C of graphite books (-8 to -2‰) and graphite flames (-24 to -10‰) depict a clear dichotomy in the same meteorite thus refuting melting.

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δ¹⁵N‰ of sinoite in FeNi is -24.9‰ in MS-17 & MS-177.

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