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## Chips off the old block: enstatite chondrites as samples of proto-Earth.

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Enstatite chondrites (EC) have a remarkable iso topic similarity to Earth, e.g.[1] but are chemical dissimilar to Earth and inner Solar system samples. Assuming a common origin, the chemical difference could be explained by re-equilibration with a gas from which O-rich solids of approximately CI com position were removed [2,3]. Charging of nebular gas by elements lost from accreting protoplanets may explain enrichment in volatiles, e.g. Na [4]. Because the (few) established EC ages precede the Moon-forming impact and post-impact Earth is iso topically similar to EC, Theia must have been simi lar too [5]. It also follows that input from the outer Solar System following EC formation at ≤10 Myr was negligible as well. We propose that: (i) proto planets grew from an isotopically uniform inner So lar System reservoir; (b) this was isolated from the outer Solar System by Jupiter; (c) variation in  $\Delta^{17}O$ was produced by differential retention of <sup>16</sup>O-poor  $^{1}{}^{26}$ Al. water in planetesimals as they were heated by Bodies that accreted quickly (e.g. Mars) retained more water with higher  $\Delta^{17}$ O; (d) EC retained their Earth-like (but not Earth-identical) O-isotopic signa tures because the residual gas was oxygen-depleted. **References:** [1] Dauphas, N. et al. 2014 EPSL 407, 96. [2] Hutson, M. & Ruzicka, A. 2000, MEPS 35, 601. [3] Gaidos, E. 2015 ApJ 804, 40. [4] Gaidos, E. 2014, Metsoc 77, 5091. [5] Herwarts, D. et al. 2014 Science 344, 1146.

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